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Particle energization in low-Mach-number fast-mode shocks in solar flares



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Abstract

During solar flares, low-Mach-number and weakly compressive fast-mode shocks can occur in magnetic reconnection outflows¹ in solar flares and are considered to be a site of electron energization responsible for the observed hard x-rays by the YOHKOH and the RHESSE measurements. In this talk we will present some recent 2D particle-in-cell simulations of collisionless low Mach/high beta shocks and the particle acceleration processes therein.^{2,3} These simulations show that collisionless plasma instabilities in the shock transition region provide the necessary



dissipation for shock formation. Both electrons and ions are found to participate in the shock-drift-acceleration (SDA), which operates differently in purely- and quasi-perpendicular shocks. SDA is most effective in accelerating electrons in quasi-perpendicular shocks. The transition energy between the thermal and non-thermal spectrum and the spectral index from the simulations are consistent with some of the X-ray spectra from RHESSI in the energy regime of $E = 40 \sim 100$ keV. At higher energy, acceleration mechanisms other than SDA are needed. This work was supported by DOE under Grant No. DE-FG02-04ER54789 and Cooperate Agreement No. DE-FC52-08NA28302, by NSF under Grant PHY-0903797, and by NSFC under Grant No. 11129503.

^[1] J. C. Workman, E. G. Blackman, and C. Ren, "Simulations reveal fast mode shocks in magnetic reconnection outflows," Phys. Plasmas 18, 092902 (2011)

^[2] J. Park, J. C. Workman, E. G. Blackman, C. Ren, and R. Siller, "Particle-in-cell simulations of particle energization from low Mach number fast mode shocks," Phys. Plasmas 19, 062904 (2012)

^[3] J. Park, C. Ren, J. C. Workman, and E. G. Blackman, "Particle-in-cell simulations of particle energization via shock drift acceleration from low Mach number quasi-perpendicular shocks in solar flares," ApJ 765, 147-157 (2013)